LUMINOSITY

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- Colliders vs. fixed-target experiments
- General definition of luminosity
- Simplest formula for Head-On (HO) collisions
- Some complications
 - Crossing angle
 - Transverse beam offset
 - Hourglass effect
- Integrated luminosity and maximization
- Pile-up, luminosity leveling and luminous region
- Summary: How to reach high luminosity?

COLLIDERS VS. FIXED-TARGET EXPERIMENTS (1/2)

 Using the relativistic equations given in Introduction, it can be seen that

$$\sqrt{s} = E_{CM} = \sqrt{m_{01}^2 c^4 + m_{02}^2 c^4 + 2(E_1 E_2 - \vec{p}_1 \cdot \vec{p}_2 c^2)}$$

For a fixed target ($\vec{p}_2 = 0$) and if we neglect the masses (i.e. if we are at sufficiently high energy)

$$\sqrt{s} = \sqrt{2 E_1 m_{02} c^2}$$

For a collider ($\vec{p}_2 = -\vec{p}_1$) and if we neglect the masses (i.e. if we are at sufficiently high energy)

$$\sqrt{s} = 2 E_1$$

COLLIDERS VS. FIXED-TARGET EXPERIMENTS (2/2)

- Numerical applications (for the available energy in the CM, i.e. to create new particles)
 - LHC (p⁺p⁺, 7 TeV / beam)
 - Collider mode => 14 TeV
 - Fixed-target mode => ~ 115 GeV (i.e. ~ 122 times less)

- LEP (e⁺e⁻, 105 GeV / beam)
 - Collider mode => 210 GeV
 - Fixed-target mode => ~ 0.3 GeV (i.e. 626 times less)

GENERAL DEFINITION OF LUMINOSITY (1/7)

 By definition, the luminosity *L* is the time-averaged integral over the interaction volume Ω of the number of reactions per unit time and volume

$$L = \frac{1}{\sigma_r T_b} \int_0^{T_b} \int_\Omega \frac{d^2 N}{dt \, d\Omega} \, dt \, d\Omega$$

Total cross section of
the reaction [m²]
Inverse of the bunch
collision frequency [s]

$$T_b^{-1} = f_b = f_{rev} M$$

GENERAL DEFINITION OF LUMINOSITY (2/7)

 The number of reactions per unit time and unit volumes satisfies the following relation associated with the Lorentz transformation of the variables => Luminosity density S

S =
$$\frac{1}{\sigma_r} \frac{d^2 N}{dt d\Omega} = N_1 N_2 \rho_1(x, y, s, t) \rho_2(x, y, s, t) M_{KLF}$$

Density of bunch 1 Density of bunch 2
 $\int d\Omega \rho_1 = 1 \int d\Omega \rho_2 = 1$ Møller Kinematic Luminosity Factor
 $M_{KLF} = \sqrt{\left(\vec{v}_1 - \vec{v}_2\right)^2 \left(-\frac{\left(\vec{v}_1 \times \vec{v}_2\right)^2}{c^2}\right)}$ Correction factor that makes S a relativistic invariant

GENERAL DEFINITION OF LUMINOSITY (3/7)

 Møller Kinematic Luminosity Factor is linked to the relative velocity between the 2 beams v₂₁ (see Introduction)

$$M_{KLF} = v_{21} \left(1 - \frac{\vec{v}_1 \cdot \vec{v}_2}{c^2} \right)$$

 \vec{v}_1 = Velocity in the Laboratory frame of all particles of bunch 1 \vec{v}_2 = Velocity in the Laboratory frame of all particles of bunch 2

m

$$= M N_1 N_2 f_{rev} M_{KLF} \int_{0}^{T_b} \int_{\Omega} \rho_1(x, y, s, t) \rho_2(x, y, s, t) dt d\Omega$$



GENERAL DEFINITION OF LUMINOSITY (5/7)

 Møller Kinematic Luminosity Factor (general case with crossing angle)

$$M_{KLF} = \sqrt{\left(\vec{v}_{1} - \vec{v}_{2}\right)^{2} - \frac{\left(\vec{v}_{1} \times \vec{v}_{2}\right)^{2}}{c^{2}}}$$



GENERAL DEFINITION OF LUMINOSITY (6/7)

$$\Rightarrow \vec{v}_1 - \vec{v}_2 = \begin{vmatrix} (v_1 - v_2) \sin \frac{\Phi}{2} \\ 0 \\ (v_1 + v_2) \cos \frac{\Phi}{2} \end{vmatrix}$$

=>

 \mathcal{V}_{1}

$$\vec{v}_1 \times \vec{v}_2 = \begin{vmatrix} 0 \\ 2 v_1 v_2 \cos \frac{\Phi}{2} \sin \frac{\Phi}{2} \\ 0 \end{vmatrix}$$

$$M_{KLF} = \sqrt{v_1^2 + v_2^2 + 2v_1v_2\cos\Phi - \frac{v_1^2v_2^2}{c^2}\sin^2\Phi}$$

$$v_{1} = \beta_{1} c$$

$$v_{2} = \beta_{2} c$$

$$\frac{M_{KLF}}{c} = \sqrt{\beta_{1}^{2} + \beta_{2}^{2} + 2\beta_{1}\beta_{2}\cos\Phi - \beta_{1}^{2}\beta_{2}^{2}\sin^{2}\Phi}$$

GENERAL DEFINITION OF LUMINOSITY (7/7)

If
$$\beta_1 = \beta_2 = \beta$$
 => $\frac{M_{KLF}}{c} = 2\beta\cos\frac{\Phi}{2}\sqrt{1-\beta^2\sin^2\frac{\Phi}{2}}$

If
$$\beta_1 = \beta_2 = \beta = 1$$
 => $\frac{M_{KLF}}{c} = 2\cos^2\frac{\Phi}{2}$

• If
$$\Phi = 0 \Rightarrow \frac{M_{KLF}}{c} = 2$$

SIMPLEST FORMULA FOR HEAD-ON COLLISIONS (1/4)

Luminosity in the absence of crossing angle (and transverse beam offset and hourglass effect)

$$L = M N_1 N_2 f_{rev} 2 \int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} \rho_1(x, y, s, -s_0) \rho_2(x, y, s, s_0) dx dy ds ds_0$$

Assuming that the densities are uncorrelated in all planes

$$\rho_{1}(x, y, s, -s_{0}) = \rho_{1x}(x)\rho_{1y}(y)\rho_{1s}(s-s_{0})$$

$$\rho_{2}(x, y, s, s_{0}) = \rho_{2x}(x)\rho_{2y}(y)\rho_{2s}(s+s_{0})$$

Assuming Gaussian distributions in all dimensions

$$\rho_{1x}(x) = \frac{1}{\sigma_{1x}\sqrt{2\pi}} e^{-\frac{x^2}{2\sigma_{1x}^2}},$$

SIMPLEST FORMULA FOR HEAD-ON COLLISIONS (2/4)

$$L = \frac{2 M N_1 N_2 f_{rev}}{(2 \pi)^3 \sigma_{lx} \sigma_{2x} \sigma_{ly} \sigma_{2y}}$$

$$\Rightarrow \int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} \int_{-\infty}^{\infty} \frac{e^{-\frac{x^2}{2\sigma_{1x}^2}} e^{-\frac{y^2}{2\sigma_{1y}^2}} e^{-\frac{(s-s_0)^2}{2\sigma_{1s}^2}} e^{-\frac{x^2}{2\sigma_{2x}^2}} e^{-\frac{y^2}{2\sigma_{2y}^2}} e^{-\frac{(s+s_0)^2}{2\sigma_{2x}^2}}}{\sigma_{1x} \sigma_{2x}} dx dy ds ds_0$$

$$\bullet \text{ Assuming } \sigma_{1s} = \sigma_{2s} = \sigma_s \implies e^{-\frac{(s-s_0)^2}{2\sigma_{1s}^2}} e^{-\frac{(s+s_0)^2}{2\sigma_{2x}^2}} = e^{-\frac{s^2}{\sigma_s^2}} e^{-\frac{s_0^2}{\sigma_s^2}}$$

$$and \iint_{-\infty}^{-\frac{s^2}{\sigma_s^2}} \frac{e^{-\frac{s_0^2}{\sigma_s^2}}}{\sigma_s^2} ds ds_0 = \pi \quad \text{(see Useful relations in Introduction)}$$

$$\Rightarrow L = \frac{M N_1 N_2 f_{rev}}{4 \pi^2 \sigma_{1x} \sigma_{2x} \sigma_{1y} \sigma_{2y}} \int_{-\infty}^{+\infty} e^{-\frac{x^2}{2\sigma_{1x}^2}} e^{-\frac{y^2}{2\sigma_{1y}^2}} e^{-\frac{x^2}{2\sigma_{2x}^2}} e^{-\frac{y^2}{2\sigma_{2y}^2}} dx dy$$

SIMPLEST FORMULA FOR HEAD-ON COLLISIONS (3/4)

• Assuming
$$\frac{\sigma_{1x} = \sigma_{2x} = \sigma_x}{\sigma_{1y} = \sigma_{2y} = \sigma_y} \text{, as} \qquad \iint \frac{e^{-\frac{x^2}{\sigma_x^2}} e^{-\frac{y^2}{\sigma_y^2}}}{\sigma_x \sigma_y} dx dy = \pi$$

one finally obtains

$$L = \frac{M N_1 N_2 f_{rev}}{4 \pi \sigma_x \sigma_y}$$
 Let's call it L_0

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• If
$$\frac{\sigma_{1x} \neq \sigma_{2x}}{\sigma_{1y} \neq \sigma_{2y}} \Rightarrow L = \frac{M N_1 N_2 f_{rev}}{2 \pi \sqrt{\sigma_{1x}^2 + \sigma_{2x}^2} \sqrt{\sigma_{1y}^2 + \sigma_{2y}^2}}$$

SIMPLEST FORMULA FOR HEAD-ON COLLISIONS (4/4)

• Assuming

$$\begin{aligned}
\sigma_{1x} &= \sigma_{2x} = \sigma_{x} \\
\sigma_{1y} &= \sigma_{2y} = \sigma_{y} \\
\sigma_{x} &= \sigma_{y} = \sigma \\
N_{1} &= N_{2} = N_{b}
\end{aligned}
= > L_{0} = \frac{M N_{b}^{2} f_{rev} \beta \gamma}{4 \pi \beta^{*} \varepsilon_{n}}$$

$$= N_{1} = N_{2} = N_{b}$$
using

$$\varepsilon_{n} = \beta \gamma \varepsilon = \beta \gamma \frac{\sigma^{2}}{\beta^{*}}$$
Normalized
transverse beam
emittance

• Numerical application for LHC => $L_0 = 1.2 \times 10^{34} \text{ cm}^{-2} \text{s}^{-1}$



SOME COMPLICATIONS: CROSSING ANGLE (2/5)

=> The following relations are thus obtained (see also Useful relations in Introduction)

$$L_{CA} = 2\cos^2 \frac{\Phi}{2} M N_1 N_2 f_{rev}$$

$$\int_{-\infty}^{+\infty}\int_{-\infty}^{+\infty}\int_{-\infty}^{+\infty}\int_{-\infty}^{+\infty}\rho_{1x}(x_{1})\rho_{1y}(y_{1})\rho_{1s}(s_{1}-s_{0})\rho_{2x}(x_{2})\rho_{2y}(y_{2})\rho_{2s}(s_{2}+s_{0})dxdydsds_{0}$$

SOME COMPLICATIONS: CROSSING ANGLE (3/5)

Assuming

$$\sigma_{1x} = \sigma_{2x} = \sigma_{x}$$
$$\sigma_{1y} = \sigma_{2y} = \sigma_{y}$$
$$\sigma_{1s} = \sigma_{2s} = \sigma_{s}$$
$$y_{1} = y_{2} = y$$

, and using the relation given in the

Useful relations in Introduction

$$\int_{-\infty}^{+\infty} e^{-\left(at^{2}+bt+c\right)} dt = \sqrt{\frac{\pi}{a}} e^{\frac{b^{2}}{4a}-c}$$

yields
$$L_{CA} = L_0 F_{CA}$$
 , with

$$F_{CA} = \frac{1}{\sqrt{1 + \left(\frac{\sigma_s}{\sigma_x} \tan \frac{\Phi}{2}\right)^2}}$$

 $\tan \Phi / 2 \sim \Phi / 2$

Numerical application for LHC
 => F_{CA} = 0.84

Elias Métral, Training-week in Accelerator Physics, Lund, Sweden, May 27-31, 2013

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SOME COMPLICATIONS: CROSSING ANGLE (4/5)



SOME COMPLICATIONS: CROSSING ANGLE (5/5)

Possibility to compensate for this geometric loss factor => Use "Crab Cavities"



=> Already used in leptons machine (KEK-B, Japan) and planned to be used for the upgrade of the LHC

SOME COMPLICATIONS: TRANSVERSE OFFEST (1/3)

Luminosity in the presence of a transverse offset (but no crossing angle)



SOME COMPLICATIONS: TRANSVERSE OFFEST (2/3)

$$L_{TO} = 2 M N_1 N_2 f_{rev}$$

$$\int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} \rho_{1x}(x_1) \rho_{1y}(y) \rho_{1s}(s-s_0) \rho_{2x}(x_2) \rho_{2y}(y) \rho_{2s}(s+s_0) dx dy ds ds_0$$

Assuming

$$\sigma_{1y} = \sigma_{2y} = \sigma_{y}$$
$$\sigma_{1s} = \sigma_{2s} = \sigma_{s}$$
$$y_{1} = y_{2} = y$$

 $\sigma_{1x} = \sigma_{2x} = \sigma_{x}$

, and using the relation given in the

Useful relations in Introduction, yields

$$L_{TO} = L_0 F_{TO}$$

with

$$F_{TO} = e^{-\left(\frac{d_1 - d_2}{2\sigma_x}\right)^2}$$

SOME COMPLICATIONS: TRANSVERSE OFFEST (3/3)



SOME COMPLICATIONS: HOURGLASS EFFECT (1/3)

- Luminosity in the presence of the Hourglass effect (and crossing angle in the x-s plane but no transverse offset)
 - Close to the IP, the beta function is given by (see course on Transverse Beam Dynamics)

$$\beta(s) = \beta^* \left[1 + \left(\frac{s}{\beta^*}\right)^2 \right]$$



SOME COMPLICATIONS: HOURGLASS EFFECT (2/3)

Following the same approach as before, yields

$$L_{CA\& HG} = \frac{\cos\frac{\Phi}{2} N_1 N_2 M f_{rev}}{4 \pi^{3/2} \sigma_s \sigma_s^* \sigma_y^*} \int_{-\infty}^{+\infty} ds \frac{e^{-s^2 \left\{ \frac{\sin^2 \Phi}{2} \left[1 + \left(\frac{s}{\beta^*} \right)^2 \right] + \frac{\cos^2 \Phi}{\sigma_s^2} \right\}}}{\left[1 + \left(\frac{s}{\beta^*} \right)^2 \right]}$$

$$\Rightarrow L_{CA\&HG} = L_{CA} F_{HG}$$

with



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SOME COMPLICATIONS: HOURGLASS EFFECT (3/3)



INTEGRATED LUMINOSITY AND MAXIMIZATION (1/4)

- Integrated luminosity $L_{int} = \int_{0}^{t} L(t) dt$
- Real figure of merit

$$L_{\rm int} \sigma_r$$
 = number of events

 Let's assume some luminosity lifetime behaviour => Exponential decay (due to intensity decay, emittance growth, etc.)

$$L(t) = L_{peak} e^{-\frac{t}{\tau_l}}$$

Luminosity lifetime

• What is the best run time t_r ?

INTEGRATED LUMINOSITY AND MAXIMIZATION (2/4)

 Let's call t_p the preparation time (time needed to put the beams in collision after the end of the previous physics fill) => Optimization of t_r and t_p gives the maximum luminosity

$$\left\langle L \right\rangle = \frac{1}{t_r + t_p} \int_{0}^{t_r + t_p} L(t) dt$$

$$\langle L \rangle = L_{peak} \tau_l \frac{1 - e^{-\frac{t_r}{\tau_l}}}{t_r + t_p}$$

INTEGRATED LUMINOSITY AND MAXIMIZATION (3/4)





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INTEGRATED LUMINOSITY AND MAXIMIZATION (4/4)

The average luminosity is maximum when

$$t_r \approx \tau_l \ln \left(1 + \sqrt{2 \frac{t_p}{\tau_l}} + \frac{t_p}{\tau_l} \right)$$

Gives ~ 15.5 h...

PILE-UP, LEVELING AND LUMINOUS REGION (1/3)

Pile-Up (PU) = Number of events / crossing for a given luminosity

$$PU = \frac{L \sigma_r}{M f_{rev}}$$

This is a limit coming from the experiments' detectors => Better to have larger number of bunches (for the same beam intensity)

 In case the pile-up is too big, luminosity leveling techniques could be used to remain at the limit => Playing with the different parameters which can reduce the luminosity (transverse beam offset, β*, etc.)



PILE-UP, LEVELING AND LUMINOUS REGION (3/3)



SUMMARY: HOW TO REACH HIGH LUMINOSITY?

High beam intensities

- High bunch intensity => More efficient (for the same beam intensity) but pile-up issue for the experiments' detectors
- High number of bunches => Less efficient but better for the pileup
- Small transverse beam sizes (small transverse emittance and beta function at the IP)
- High energy
- Small crossing angle
- Small transverse offset
- Short bunches