

Transverse Oscillations for Coasting Beams with Strong Space Charge

A. Burov

CERN Talk, Apr 4, 2012

Möhl – Schönauer Equation

- In 1974, D. Möhl and H. Schönauer suggested to describe coasting beam oscillations by the following equation:

$$\frac{d^2 x_i}{dt^2} + \underbrace{\Omega_i^2 Q_i^2}_{\text{lattice}} x_i + 2\Omega_0^2 Q_0 \left[\underbrace{Q_c}_{\text{wake}} \bar{x} + \underbrace{Q_{sc}}_{\text{space charge}} (x_i - \bar{x}) \right] = 0.$$

$$Q_{sc} = Q_{sc}(\mathbf{J}_i); \quad Q_{sc}(0) = -\frac{N r_0}{4\pi\beta^2 \gamma^3 \epsilon_{rms}}.$$

↑
action

Coherent tune shift, in a frequency domain:

$$Q_c = -i \frac{N r_0 \beta_x}{\gamma C} \frac{Z_x}{Z_0}; \quad Z_0 = \frac{4\pi}{c} = 377 \text{ Ohm}$$

Space charge term

$$\frac{d^2 x_i}{dt^2} + \Omega_i^2 Q_i^2 x_i + 2\Omega_0^2 Q_0 \left[Q_c \bar{x} + Q_{sc} (x_i - \bar{x}) \right] = 0.$$

- The only non-trivial term in this equation is the space charge force:

Space charge term

$$\frac{d^2 x_i}{dt^2} + \Omega_i^2 Q_i^2 x_i + 2\Omega_0^2 Q_0 \left[Q_c \bar{x} + Q_{sc} (x_i - \bar{x}) \right] = 0.$$

- The only non-trivial term in this equation is the space charge force:

$$Q_{sc} (x_i - \bar{x})$$

Space charge term

$$\frac{d^2 x_i}{dt^2} + \Omega_i^2 Q_i^2 x_i + 2\Omega_0^2 Q_0 \left[Q_c \bar{x} + Q_{sc} (x_i - \bar{x}) \right] = 0.$$

- The only non-trivial term in this equation is the space charge force:

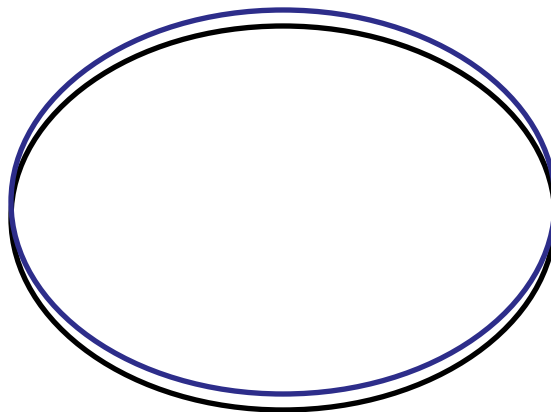
$$Q_{sc} (x_i - \bar{x})$$

- Why non-linear space charge forces are described by a linear term? Does it mean that it is valid only for K-V (constant density profile) distribution?
- No, it is valid for any beam profile. The reason is following:

Comments to Möhl-Schönauer Equation (MSE)

Space charge term: $Q_{sc}(\mathbf{J}_{\perp})(x_i - \bar{x})$

- The single-particle motion consists of 2 parts:
 - free oscillations (typically with beam size amplitude, or $J_{\perp} \simeq \varepsilon_{\text{rms}}$)
 - driven by the coherent offset oscillations (much smaller than the beam size).
- The equation describes the driven oscillations only, so it results from linearization of the original non-linear space charge term over infinitesimally small coherent motion, and averaging over the betatron phases of the free incoherent oscillations.



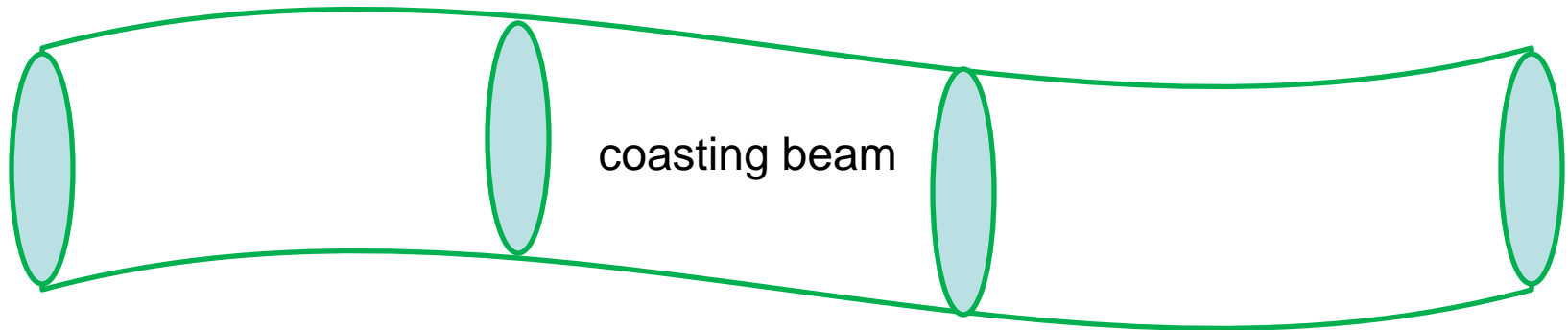
MSE as a unique choice

- This equation is a unique possibility for
 - A linear equation (driven oscillations are small!)
 - With time-independent coefficients (more on that below...)
 - With the given tunes and incoherent SC tune shifts $Q_{sc}(\mathbf{J}_\perp)$
 - Space charge force $\propto (x_i - \bar{x})$

Time-independent $Q_{sc}(\mathbf{J}) = \text{rigid-slice approximation}$

$$\frac{d^2 x_i}{dt^2} + \underbrace{\Omega_i^2 Q_i^2}_{\text{lattice}} x_i + 2\Omega_0^2 Q_0 \left[\underbrace{Q_c}_{\text{wake}} \bar{x} + \underbrace{Q_{sc}}_{\text{space charge}} (x_i - \bar{x}) \right] = 0.$$

- The assumption for the rigid slice approximation is that a core of any beam slice moves as a whole, so there is no inner motion in it.



- This is a good approximation, when the space charge is strong enough:

$$|Q_{sc}| \gg \Delta Q_i.$$

- Otherwise, the beam shape oscillates as well, and MSE is not necessarily valid.

Coasting beam: Landau damping (Burov, Lebedev, 2008):

$$\Lambda = -\pi \langle \Delta Q_{sep} \rangle \int \Delta Q_{sep} f_x J_x \delta(\Delta Q_l + Q_{sc} - \text{Re } \nu_c) d\Gamma, \quad f_x \equiv \partial f / \partial J_x ;$$

$$\Delta Q_{sep} \equiv \text{Re } \Delta Q_c - Q_{sc}(J_x, J_y), \quad \langle \Delta Q_{sep} \rangle \equiv - \left(\int \frac{f_x J_x d\Gamma}{\Delta Q_{sep}} \right)^{-1} ;$$

$$\text{Re } \nu_c = \text{Re } \Delta Q_c + \delta Q^{(1)} + \delta Q^{(2)},$$

$$\text{Im } \nu_c = \text{Im } \Delta Q_c - \Lambda$$

$$\delta Q^{(1)} = - \langle \Delta Q_{sep} \rangle \int \frac{\Delta Q_l f_x J_x d\Gamma}{\Delta Q_{sep}},$$

$$\delta Q^{(2)} = - \langle \Delta Q_{sep} \rangle \int \frac{\Delta Q_l^2 f_x J_x d\Gamma}{\Delta Q_{sep}^2}.$$

Landau Damping

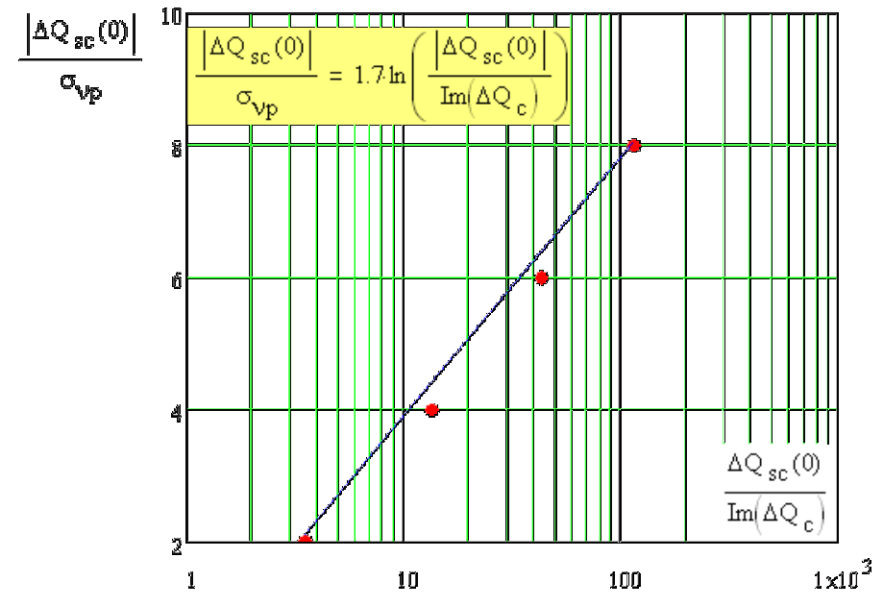
- Landau damping is proportional to the phase space density of the resonance particle:

$$\Lambda = -\pi \langle \Delta Q_{sep} \rangle \int \Delta Q_{sep} f_x J_x \delta(\Delta Q_l + Q_{sc} - \text{Re} v_c) d\Gamma$$

- Thus, with strong space charge, it is determined by the distribution tails.
- Typically, the far tails are neither predictable nor measurable, and not even reproducible in many cases. That is why, at strong space charge, the stability thresholds are predictable and reproducible with a moderate accuracy. In the RR it was 20-30% in the intensity.

Coasting Beam Thresholds (Burov, Lebedev, 2008)

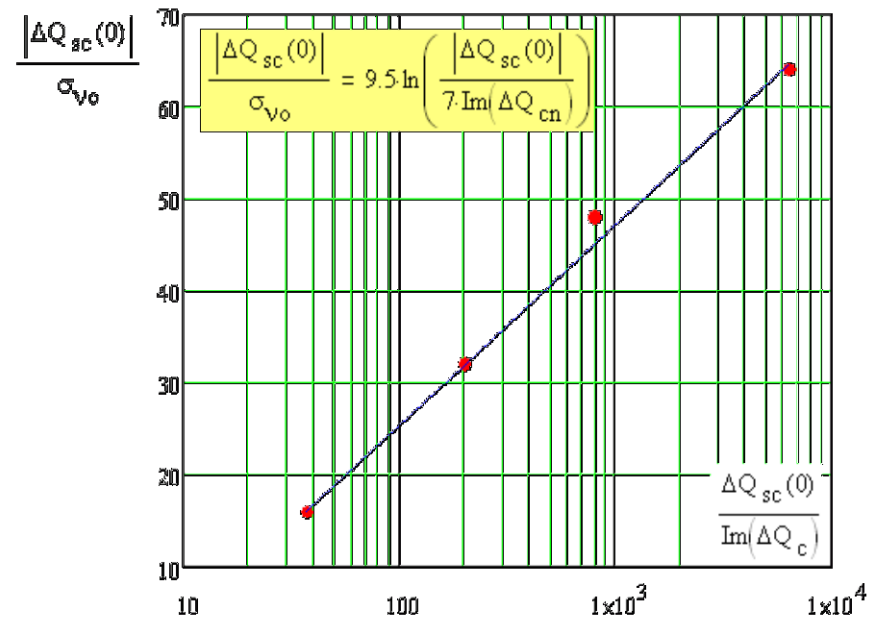
- Thresholds are determined by $\Lambda = \text{Im} \Delta Q_c$. For a round Gaussian beam:



Chromatic threshold

$$\sigma_{vp} \equiv |\xi - n\eta| \left(\frac{\delta p}{p} \right)_{\text{rms}}$$

$$\Delta Q_l = (\xi - n\eta) \hat{p} \equiv \sigma_{vp} \hat{p} / \sigma_p$$



Octupole threshold

$$\Delta Q_i = \sigma_{vo} (J_x + J_y) / 2\varepsilon > 0;$$

$$\langle J_x \rangle = \langle J_y \rangle = \varepsilon$$

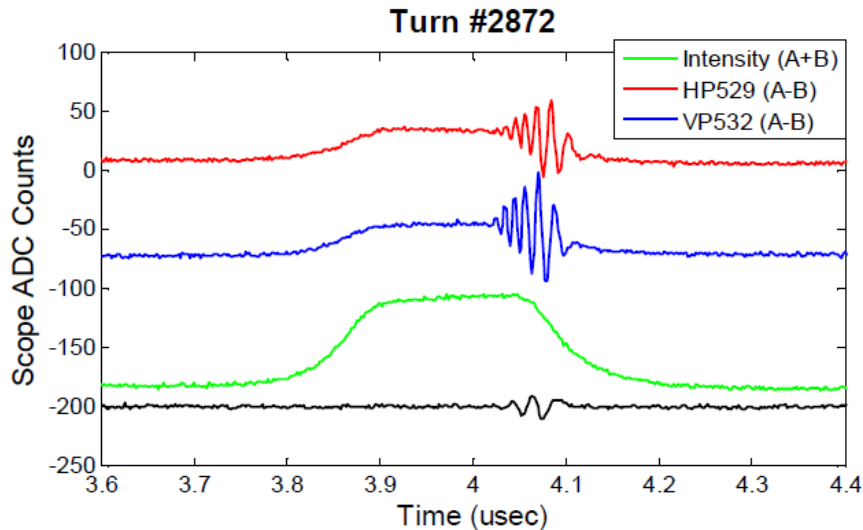
Direct application to bunched beams

- Coasting beam results sometimes can be applied to bunched beams.
 - For short-wavelength and fast oscillations (much shorter than the bunch length, much faster than the synchrotron frequency) – the so-called microwave instability;
 - For multi-bunched beam, at zero chromaticity and long wavelength (much longer than the bunch spacing)

Antiproton instabilities at Fermilab Recycler Ring

Typical instability snapshot (L. Prost et al, 2011,

<http://lss.fnal.gov/archive/test-tm/2000/fermilab-tm-2498-ad.pdf>):



- Instability was always at the tail
- Frequency 70 MHz – by the damper
- Threshold prediction : $\pm 30\%$